

we write  $G(x - y)$  in terms of the complete set of eigenfunctions  $u_k$  as

$$G(x - y) = \sum_{k=1}^{\infty} \frac{u_k(x)u_k(y)}{\lambda_k} \quad (6.410)$$

so that the action  $Lu_k = \lambda_k \rho u_k$  turns  $G$  into

$$LG(x - y) = \sum_{k=1}^{\infty} \frac{L u_k(x)u_k(y)}{\lambda_k} = \sum_{k=1}^{\infty} \rho(x) u_k(x) u_k(y) = \delta(x - y) \quad (6.411)$$

our  $\alpha = 1$  series expansion (6.374) of the delta function.

### 6.39 Green's Functions in One Dimension

In one dimension, we can explicitly solve the inhomogeneous ordinary differential equation  $L f(x) = g(x)$  in which

$$L = -\frac{d}{dx} \left( p(x) \frac{d}{dx} \right) + q(x) \quad (6.412)$$

is formally self adjoint. We'll build a Green's function from two solutions  $u$  and  $v$  of the homogeneous equation  $Lu(x) = Lv(x) = 0$  as

$$G(x, y) = \frac{1}{A} [\theta(x - y)u(y)v(x) + \theta(y - x)u(x)v(y)] \quad (6.413)$$

in which  $\theta(x) = (x + |x|)/(2|x|)$  is **the Heaviside step function** (Oliver Heaviside 1850–1925), and  $A$  is a constant which we'll presently identify. We'll show that the expression

$$f(x) = \int_a^b G(x, y) g(y) dy = \frac{v(x)}{A} \int_a^x u(y) g(y) dy + \frac{u(x)}{A} \int_x^b v(y) g(y) dy$$

solves our inhomogeneous equation. Differentiating, we find after a cancellation

$$f'(x) = \frac{v'(x)}{A} \int_a^x u(y) g(y) dy + \frac{u'(x)}{A} \int_x^b v(y) g(y) dy. \quad (6.414)$$

Differentiating again, we have

$$\begin{aligned}
 f''(x) &= \frac{v''(x)}{A} \int_a^x u(y) g(y) dy + \frac{u''(x)}{A} \int_x^b v(y) g(y) dy \\
 &\quad + \frac{v'(x)u(x)g(x)}{A} - \frac{u'(x)v(x)g(x)}{A} \\
 &= \frac{v''(x)}{A} \int_a^x u(y) g(y) dy + \frac{u''(x)}{A} \int_x^b v(y) g(y) dy \\
 &\quad + \frac{W(x)}{A} g(x)
 \end{aligned} \tag{6.415}$$

in which  $W(x)$  is the wronskian  $W(x) = u(x)v'(x) - u'(x)v(x)$ . The result (6.266) for the wronskian of two linearly independent solutions of a self-adjoint homogeneous ODE gives us  $W(x) = W(x_0)p(x_0)/p(x)$ . We set the constant  $A = -W(x_0)p(x_0)$  so that the last term in (6.415) is  $-g(x)/p(x)$ . It follows that

$$Lf(x) = \frac{[Lv(x)]}{A} \int_a^x u(y) g(y) dy + \frac{[Lu(x)]}{A} \int_x^b v(y) g(y) dy + g(x) = g(x). \tag{6.416}$$

But  $Lu(x) = Lv(x) = 0$ , so we see that  $f$  satisfies our inhomogeneous equation  $Lf(x) = g(x)$ .

## 6.40 Nonlinear Differential Equations

The field of nonlinear differential equations is too vast to cover here, but we may hint at some of its features by considering some examples from cosmology and particle physics.

The Friedmann equations of general relativity (11.410 & 11.412) for the scale factor  $a(t)$  of a homogeneous, isotropic universe are

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + 3p) \quad \text{and} \quad \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3}\rho - \frac{k}{a^2} \tag{6.417}$$

in which  $k$  respectively is 1, 0, and  $-1$  for closed, flat, and open geometries. (The scale factor  $a(t)$  tells how much space has expanded or contracted by the time  $t$ .) These equations become more tractable when the energy density  $\rho$  is due to a single constituent whose pressure  $p$  is related to it by an equation of state  $p = w\rho$ . Conservation of energy  $\dot{\rho} = -3(\rho + p)/a$  (11.426–11.431) then ensures (exercise 6.30) that the product  $\rho a^{3(1+w)}$  is